





Status update on the analysis of ⁸⁰Se(n,γ)

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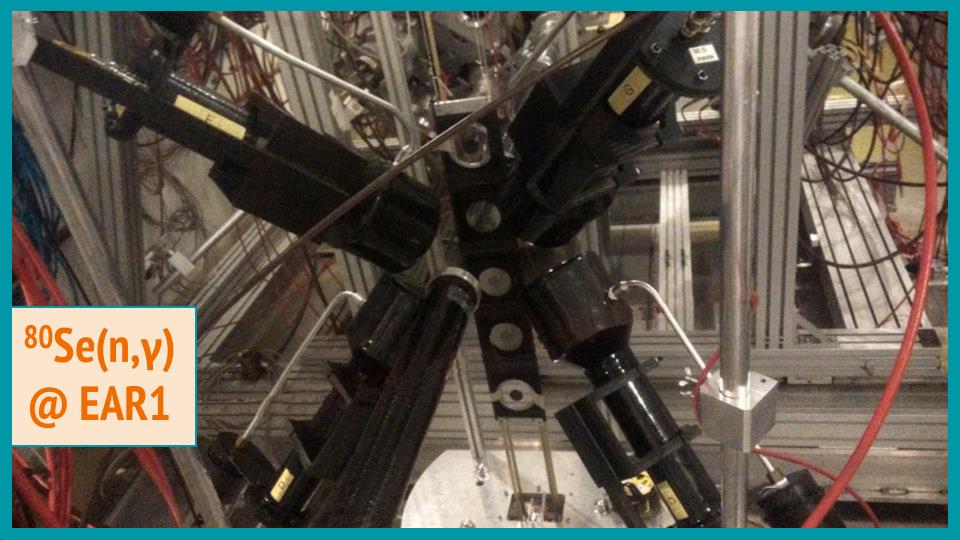
Carlos Guerrero, Mª Ángeles Millán, Mª Teresa Rodríguez (US)

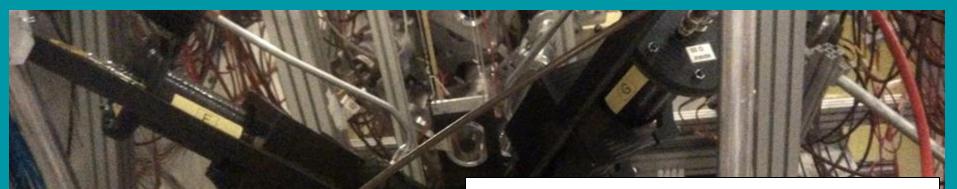
n_TOF local team and the n_TOF Collaboration







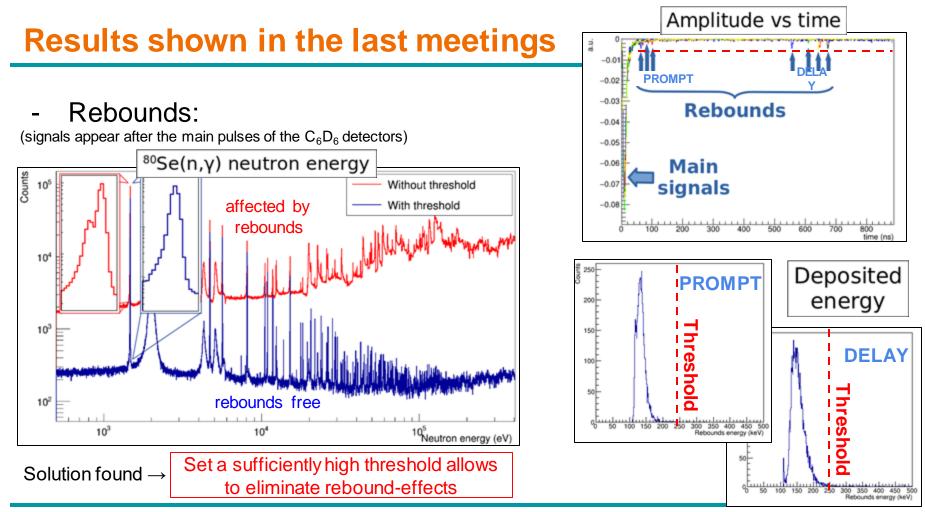




⁸⁰Se(n,γ) **@ EAR1**

Shown previously

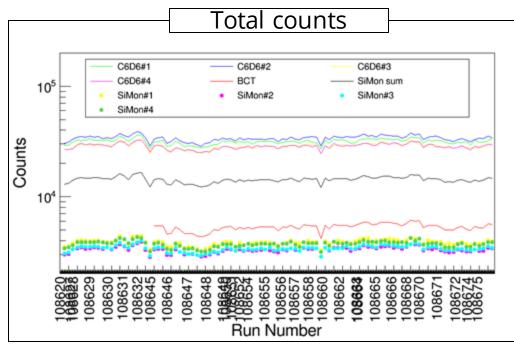
- Rebounds study (<u>Granada 18</u>).
- Count rate consistency study (<u>CERN 19</u>).
- Gain stability study (<u>Prague 19</u>).
- G4 simulation and detector calibration (Prague 19).
- Preliminary weighting function calculation (Prague 19).



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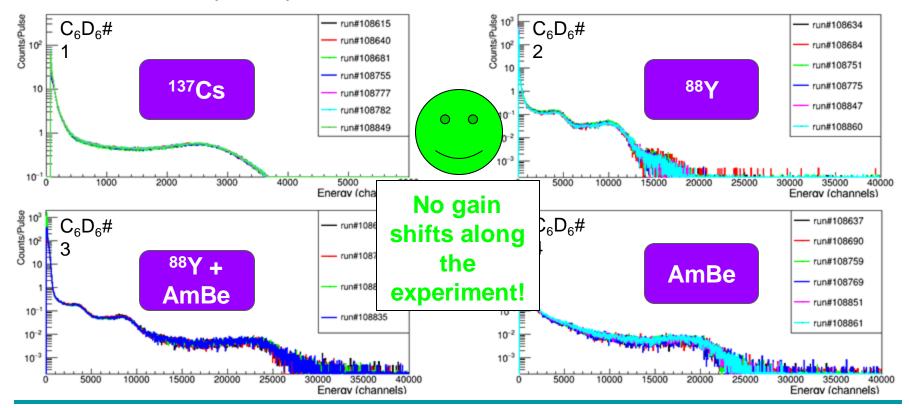
- Count rate consistency study

(check the consistency of the C₆D₆ count rates along the experiment)



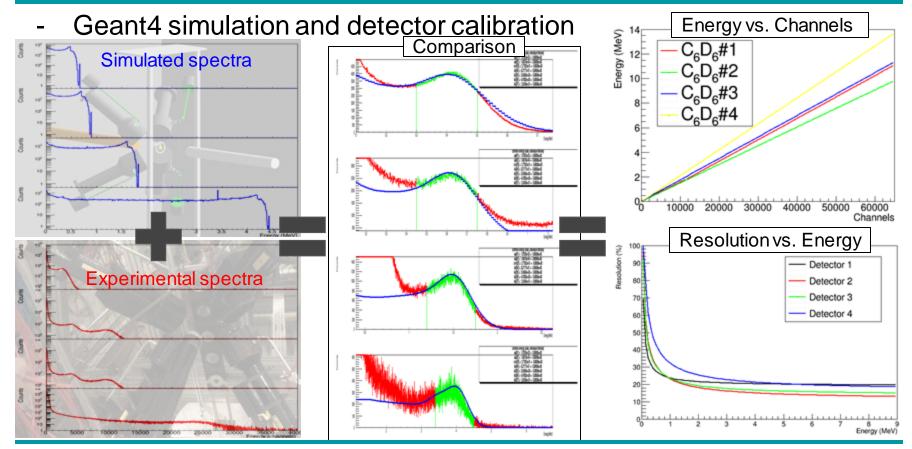
ent)						
All the statistic has been verified						
Sample	%					
Au	100					
Empty	100					
Empty + Filters	100					
Se Thick	100					
Se Thick + F	100					
Se Thin	100					

- Gain stability study (check the gain stability of all C6D6)



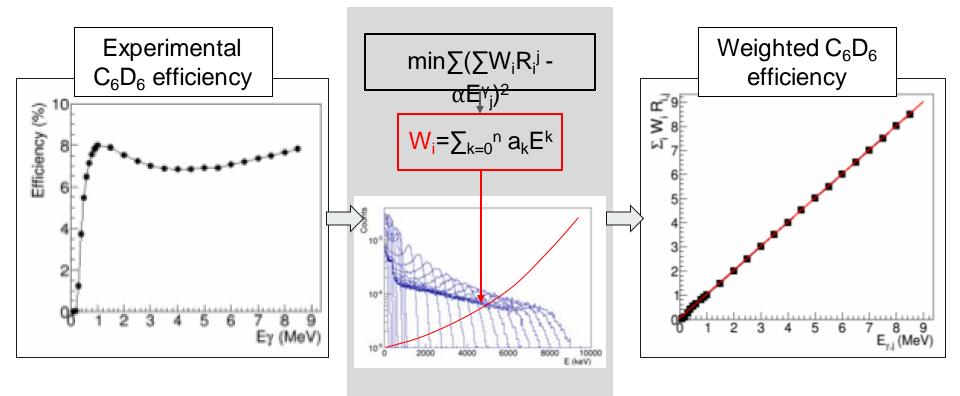
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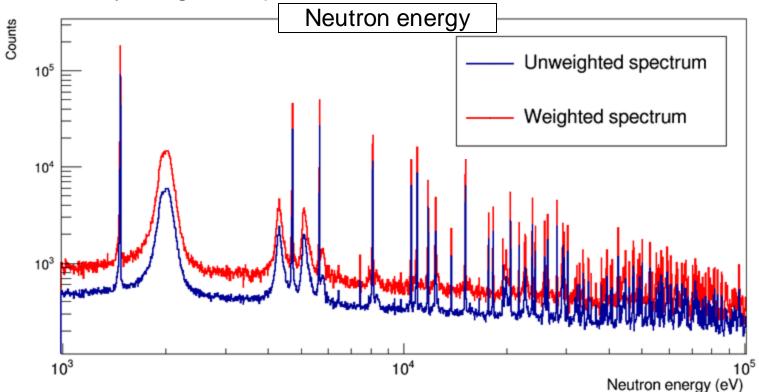


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- Weighting function calculation

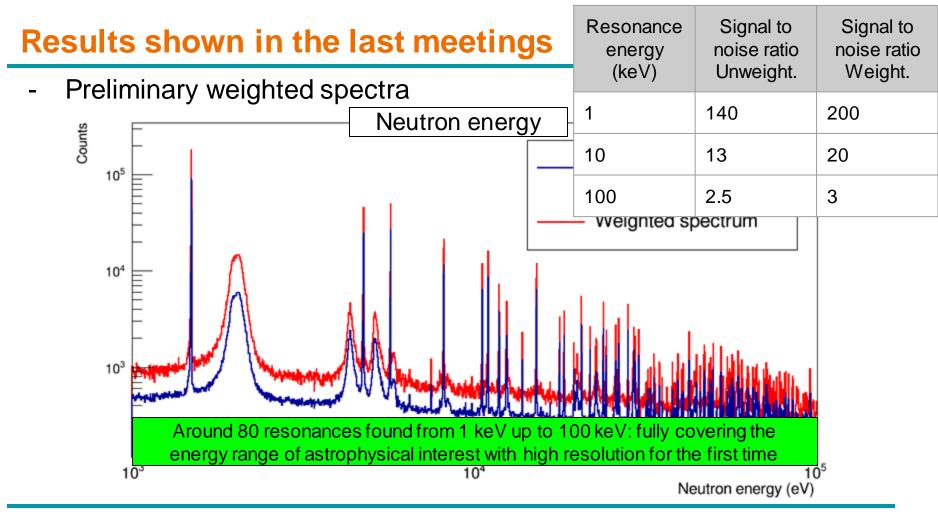


- Preliminary weighted spectra



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Ongoing work

Cascades simulation:

An event generator for simulations of complex β -decay experiments D. Jordan^{a,*}, A. Algora^{a,b}, J.L. Tain^a

Why? Calculate uncertainties in the weighting function and threshold correction factors (particularly important for the threshold used to reject rebounds).

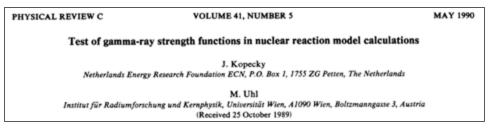
How? Create cascades with CAPTUGENS, and simulate them using Geant4.

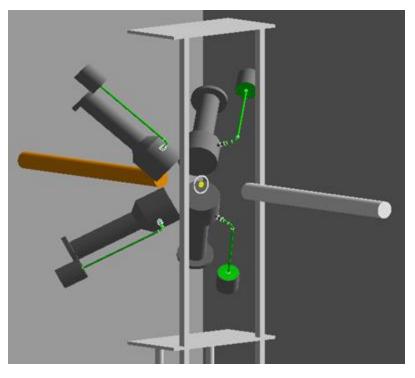
- Check methodology with Au cascade.
- Obtain the ⁸⁰Se cascade:
 - Select model to parametrize statistical part of nuclei.
 - Calculate Photon Strength Function.
 - Simulate cascade and check with experimental spectra.

- Generating the cascade for 4.9 eV resonance
- Photon Strength function:

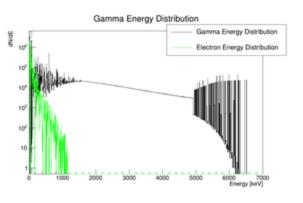
	E1		M1	E2
E (MeV)	13.72	5.8	7.05	10.81
Γ (MeV)	4.61	1.5	4.00	3.73
σ_0 (mb)	541.0	6.0	1.12	5.03

E1: two giant resonances parametrized with generalized lorentzians. M1, E2: parametrized with lorentzians.



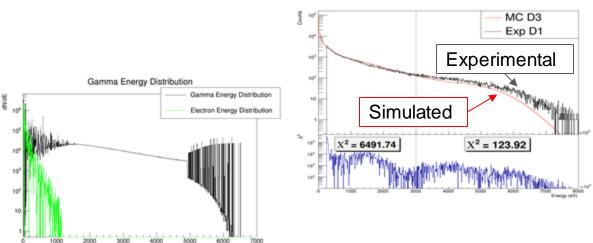


- Simulated cascade for 4.9 eV resonance





Simulated cascade for 4.9 eV resonance





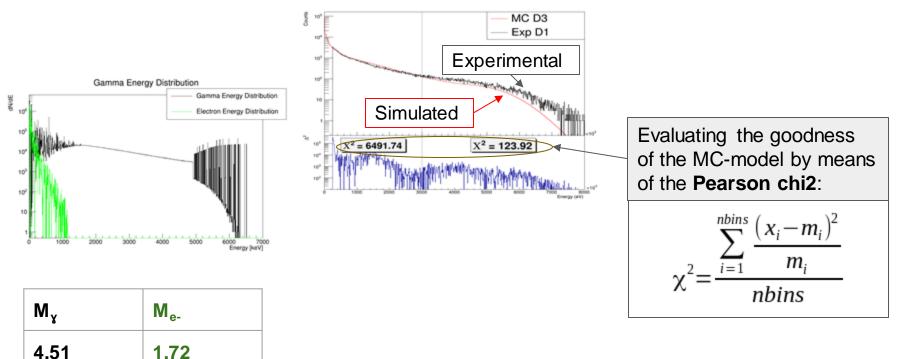
5000

6000 7000 Energy [keV]

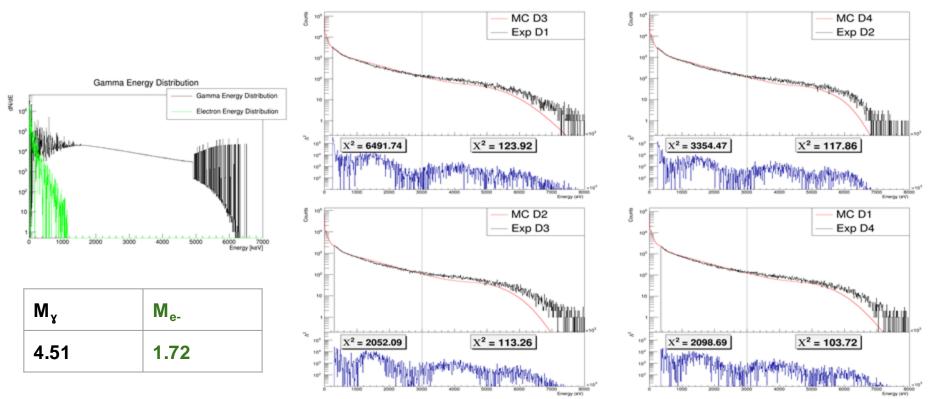
1000

2000

- Simulated cascade for 4.9 eV resonance

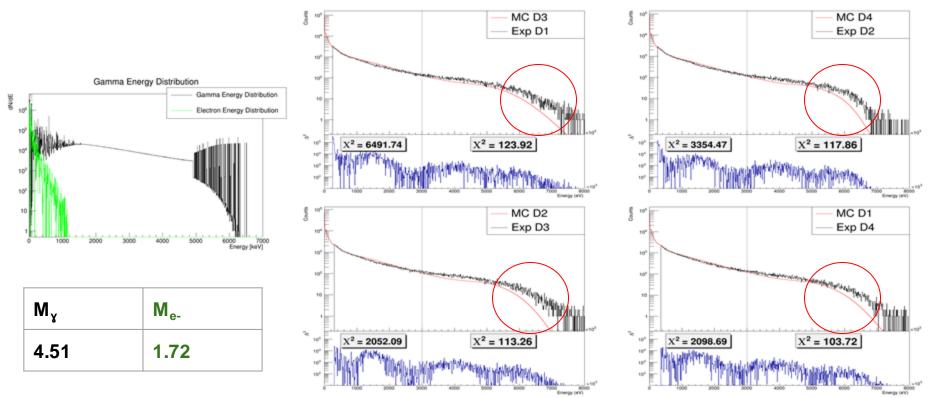


- Simulated cascade for 4.9 eV resonance



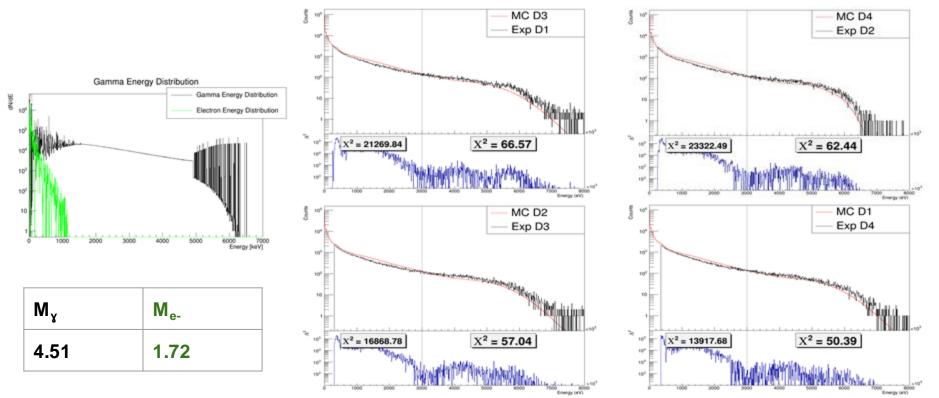
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- Simulated cascade for 4.9 eV resonance



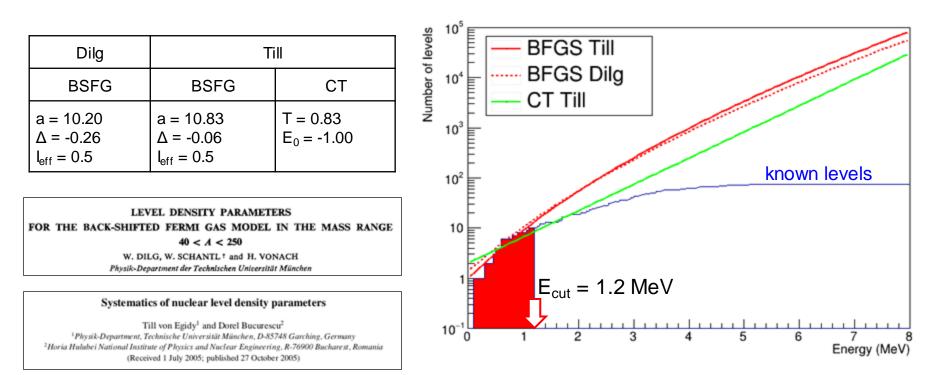
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- Simulated cascade for 4.9 eV resonance as a calibration point



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- Generating the cascade: known levels vs. statistical part.



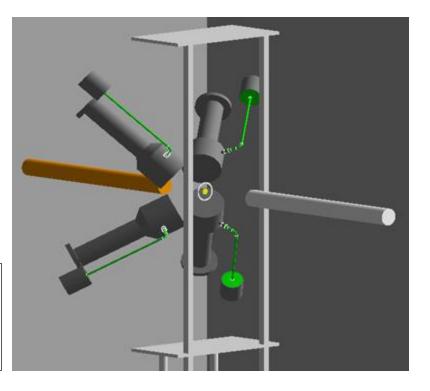
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- Generating the cascade: Photon Strength Function calculation.

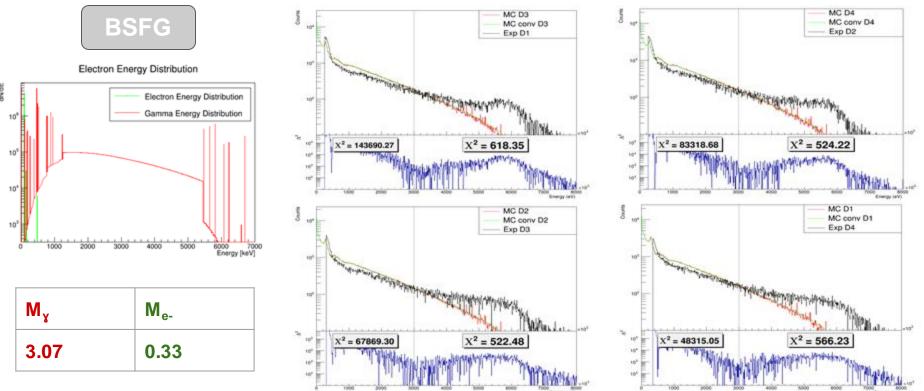
	E1		M1	E2
E (MeV)	15.44	17.89	9.48	14.56
Г (MeV)	4.85	6.42	4.00	4.41
σ ₀ (mb)	51.16	102.31	12.46	0.14

RIPL – Reference Input Parameter Library for Calculation of Nuclear Reactions and Nuclear Data Evaluations

R. Capote,^{1*} M. Herman,^{1,2} P. Obložinský,^{1,2} P.G. Young,³ S. Goriely,⁴ T. Belgya,⁵ A.V. Ignatyuk,⁶ A.J. Koning,⁷ S. Hilaire,⁸ V.A. Plujko,⁹ M. Avrigeanu,¹⁰ O. Bersillon,⁸ M.B. Chadwick,³ T. Fukahori,¹¹ Zhigang Ge,¹² Yinlu Han,¹² S. Kailas,¹³ J. Kopecky,¹⁴ V.M. Maslov,¹⁵ G. Reffo,¹⁶ M. Sin,¹⁷ E.Sh. Soukhovitskii,¹⁵ P. Talou³

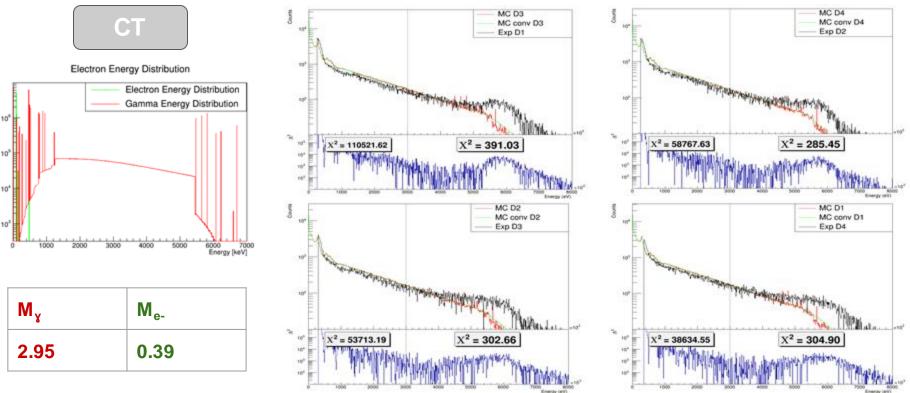


- Simulated cascade for 1.95 keV resonance: BSFG level-density.



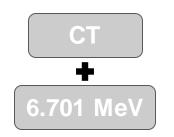
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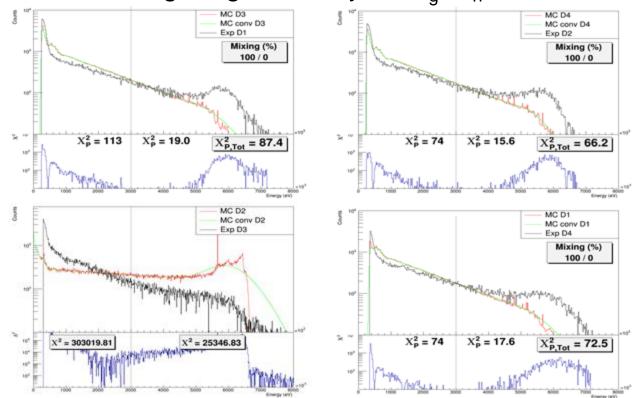
- Simulated cascade for 1.95 keV resonance: CT level-density.



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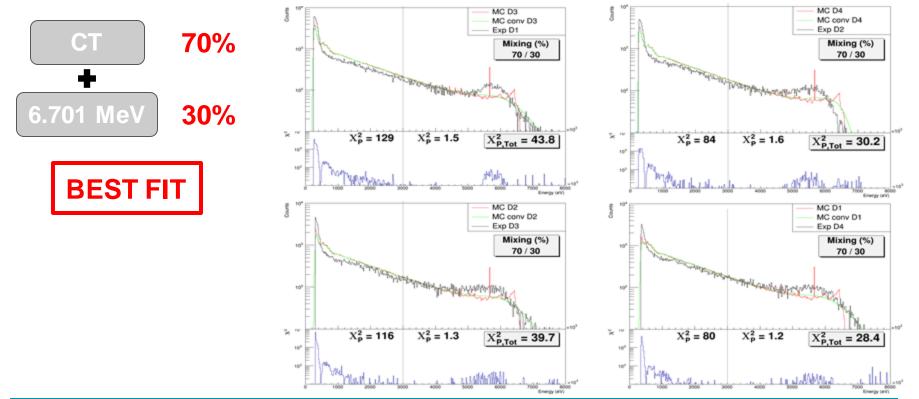
- Mixing cascade with monoenergetic gamma-rays of $E_a = S_n$





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- Mixing cascade with monoenergetic gamma-rays of Eg=S_n



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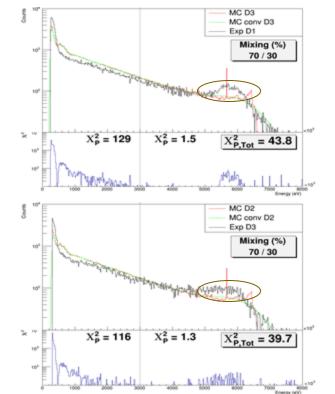
- Mixing cascade with monoenergetic gamma-rays of Eg=S_n

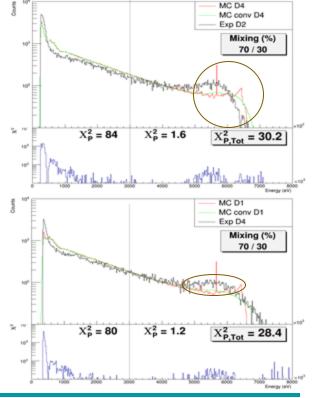


But still with aspects to improve:

- High energy "bump" does not fully reproduce.
- Calibration problem (?) in Detector#2.

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Summary and outlook

Analysis of the experiment is now ongoing.

Until now, we have studied:

- Effects of rebounds.
- The count rate stability for C_6D_6 .
- The gain stability of the C₆D₆.
- Weighting function.
- Geant4 simulations.

Study currently ongoing:

- Improve modeling of the cascades.
- Determine the WF accuracy.
- Determine corrections for the experimental effects of threshold

Next steps:

- Determine the experimental capture yield
- R-Matrix analysis with SAMMY.
- Astrophysical interpretations of the results.